THE VISUAL COMPLEXITY OF CORONAL MASS EJECTIONS

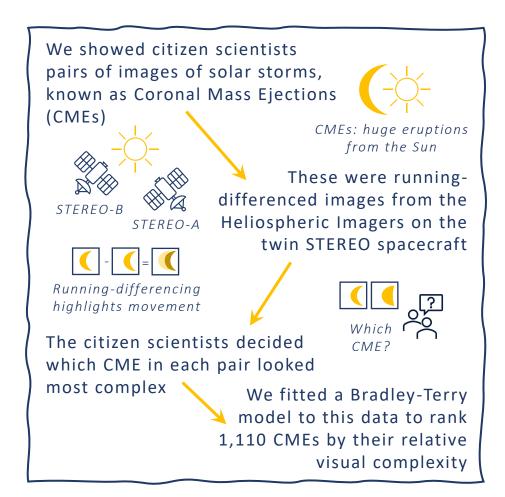
SCIENCE MUSEUM A research project by:

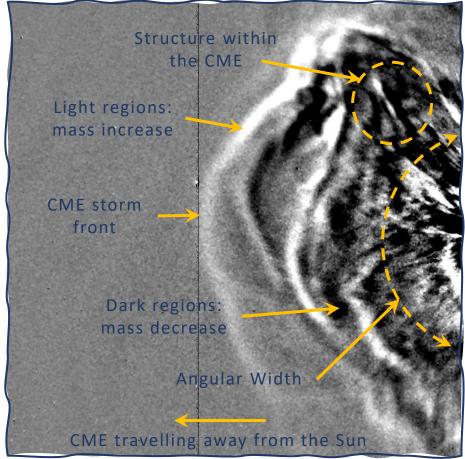
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ZOONIVERSE

E. Baeten and 4,027 other Zooniverse citizen scientists



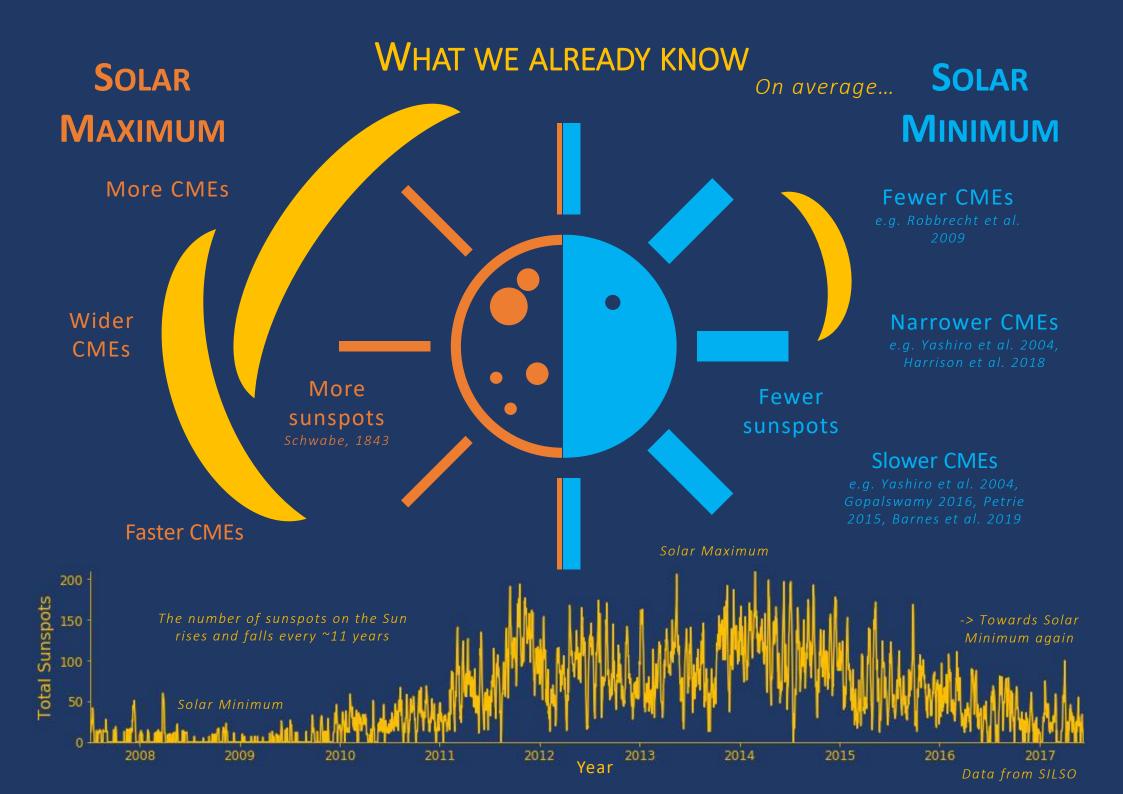


The methods we used in our study.

Our citizen science project can be viewed here:

Protect our Planet from Solar Storms

An example labelled running-differenced image of a Coronal Mass Ejection (CME) taken by the Heliospheric (white-light wide-angle) Imagers on the STEREO spacecraft



SOLAR

CONCLUSIONS FROM OUR RESEARCH

On average...

SOLAR

MINIMUM

MAXIMUM

CMEs appear more complex

image: +2.02

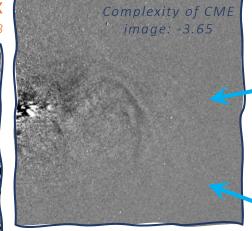
More "bright" pixels

Complexity is correlated with the number of pixel values in the top (brightest) 12.5% of the colour map.

Correlation coefficients: All STEREO-A CMEs: +0.63 All STEREO-B CMEs: +0.64

Wider

Complexity of CME



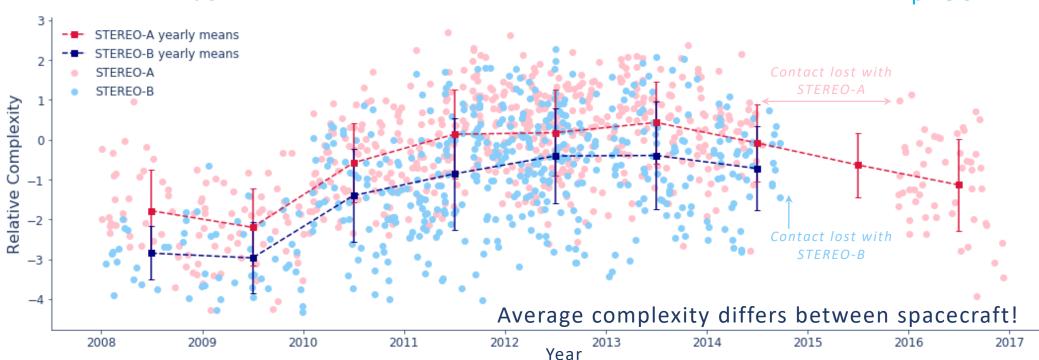
CMEs appear less complex STEREO-A CMEs averaged -2.2 in 2009

Narrower

Complexity is correlated with angular width, as measured in the HELCATS catalogue.

Correlation coefficients: All STEREO-A CMEs: +0.60 All STEREO-B CMEs: +0.43

Fewer "bright" pixels



WHAT DOES THIS MEAN?

The appearance of CMEs is changing through the solar cycle. This suggests further research is necessary to:

Quantitatively determine which CME features are changing

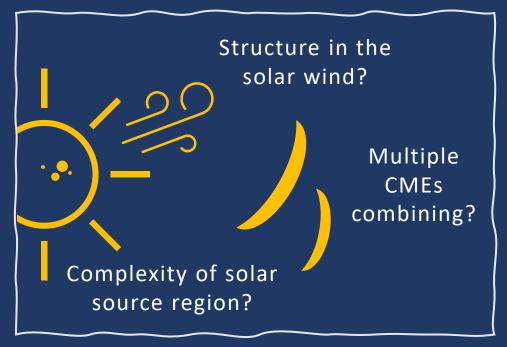
patterns
size variety direction
quality
ridges

Messiness
waves shading
brightness

storm-fronts

The qualitative CME characteristics which the citizen scientists used to describe a complex CME

Investigate what is causing the CMEs to appear differently



Cartoon showing possible reasons for the appearance of CMEs to change over the solar cycle

Our results suggest that there is some predictability in the structure of coronal mass ejections, which may help to improve future space weather forecasts.